

MWA Real-Time Subsystem

MWA post-correlation calibration & imaging sys.

Short-term picture

- C-code development / testing in Cambridge
- use simulated data now, and 32T data later

Long-term picture

- parallelized RTC implementation: 16 GB s⁻¹ thru O(100) pipes

- construct dirty images every 8^s, remove ionospheric distortion

- provide real-time calibration to other subsystems (B < 1.5 km)

Greenhill: Harvard-Smithsonian CfA

RTS Manpower & Resources

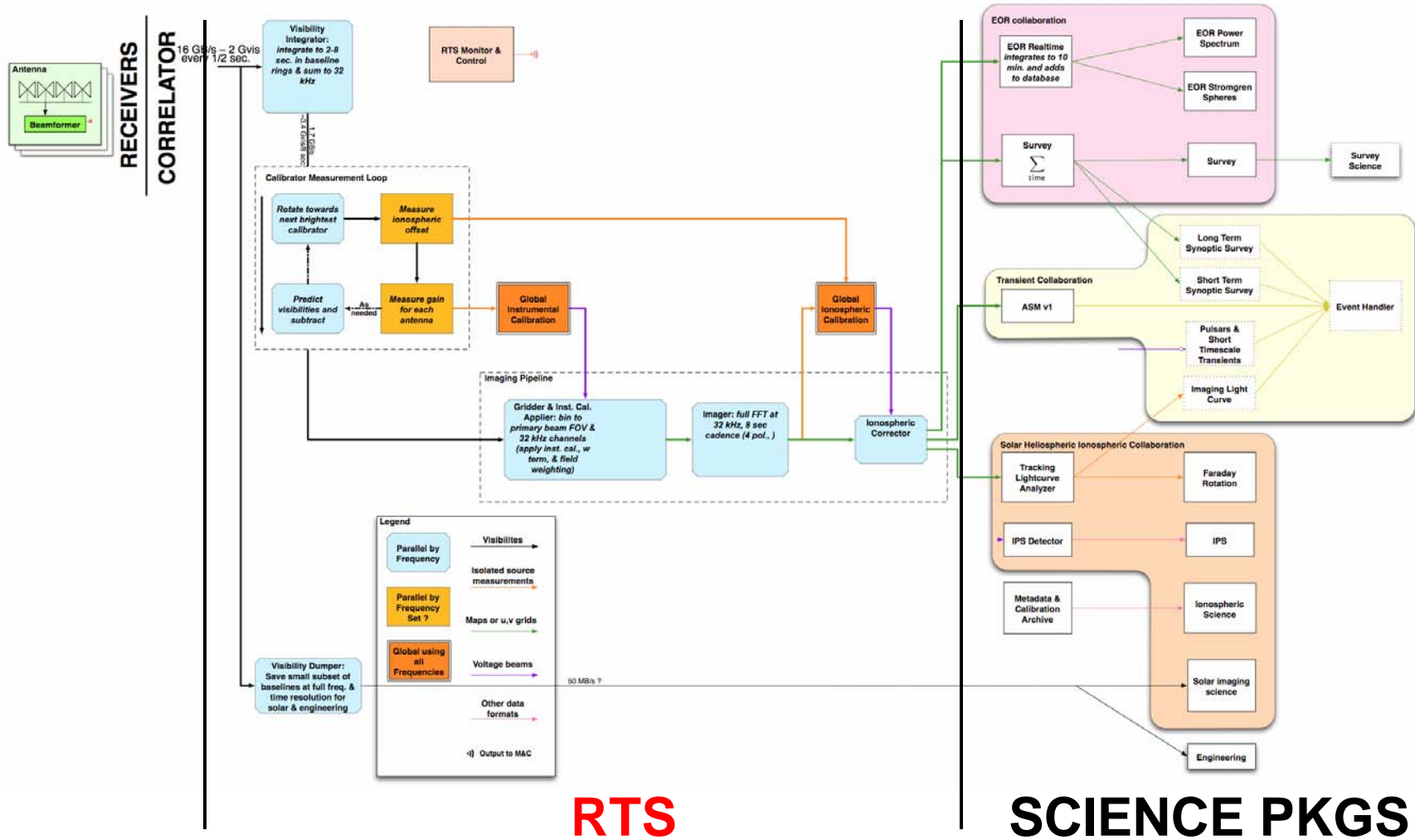
- Group kept small, geographically proximate
- Overseen by Greenhill, reporting to Sault (RT System)
 - Daniel Mitchell (SAO): **full-time**
 - Randall Wayth (SAO): **full-time**
 - Shepherd Doeleman (Haystack): *part-time*
 - Miguel Morales (MIT): *part-time*
 - Justin Kasper (MIT): *part-time*
 - Bob Sault on-site with local MWA groups 10/06 - 01/07
 - Steve Ord (SAO): new start Jul. 1, *part-time* on RTS
 - Technical requirements TBD for (1-2) open slots
- **Harvard IIC collaboration** ⇒ 200+ core devel. cluster

RTS Development Priorities

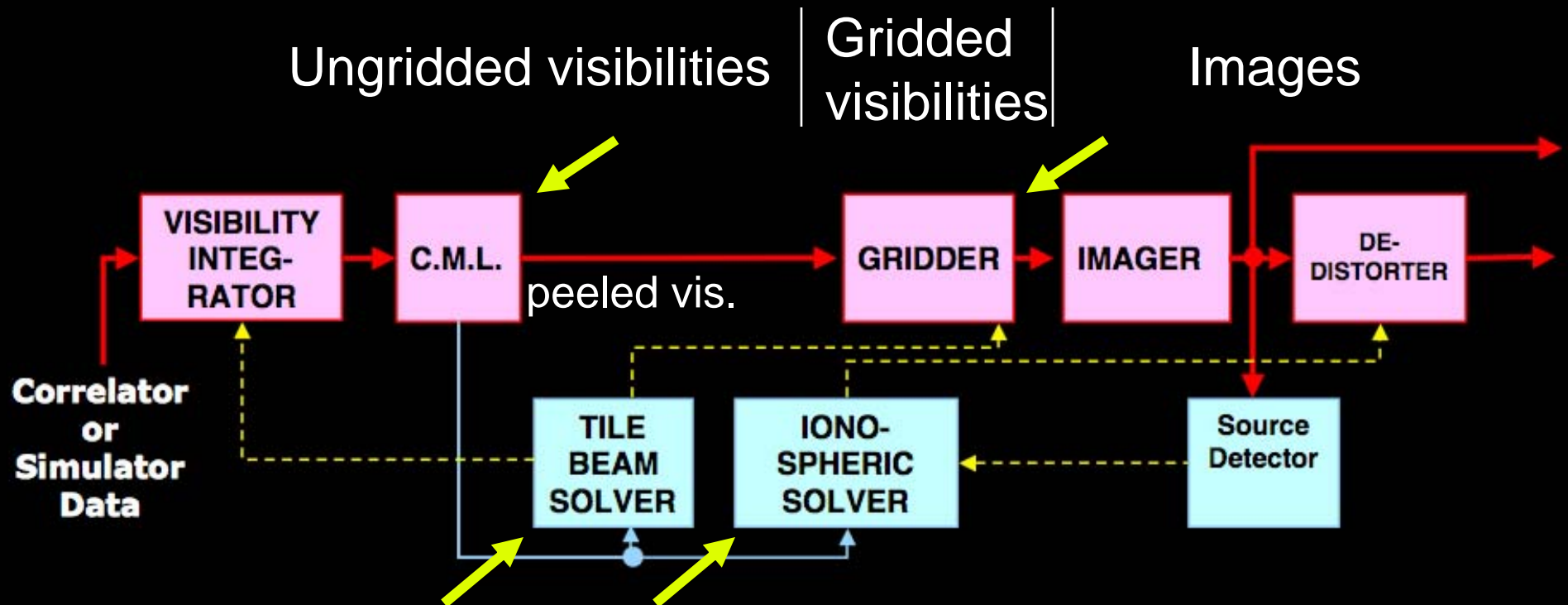
(following 03/07 TAC review)

- **Accelerated assembly of formal test harness**
 - "right answer" for RTS design cannot be anticipated
 - nuts and bolts testing required; sooner the better
 - sky, ionosphere, and MWA configuration present "challenges"
 - extensive use of simulations for test & evaluation
 - obtain benchmarks; scale; assess performance; rework; repeat
- **Drop standalone, individual exploratory codes**
 - test-harness to run on desktops and IIC cluster; **tight version ctl.**
 - built from scratch; C code
 - maintain flexibility in light of uncertain RTC configuration
- **Obtain realistic sims via MAPS**: dipoles, ionosphere...

MWA Geography

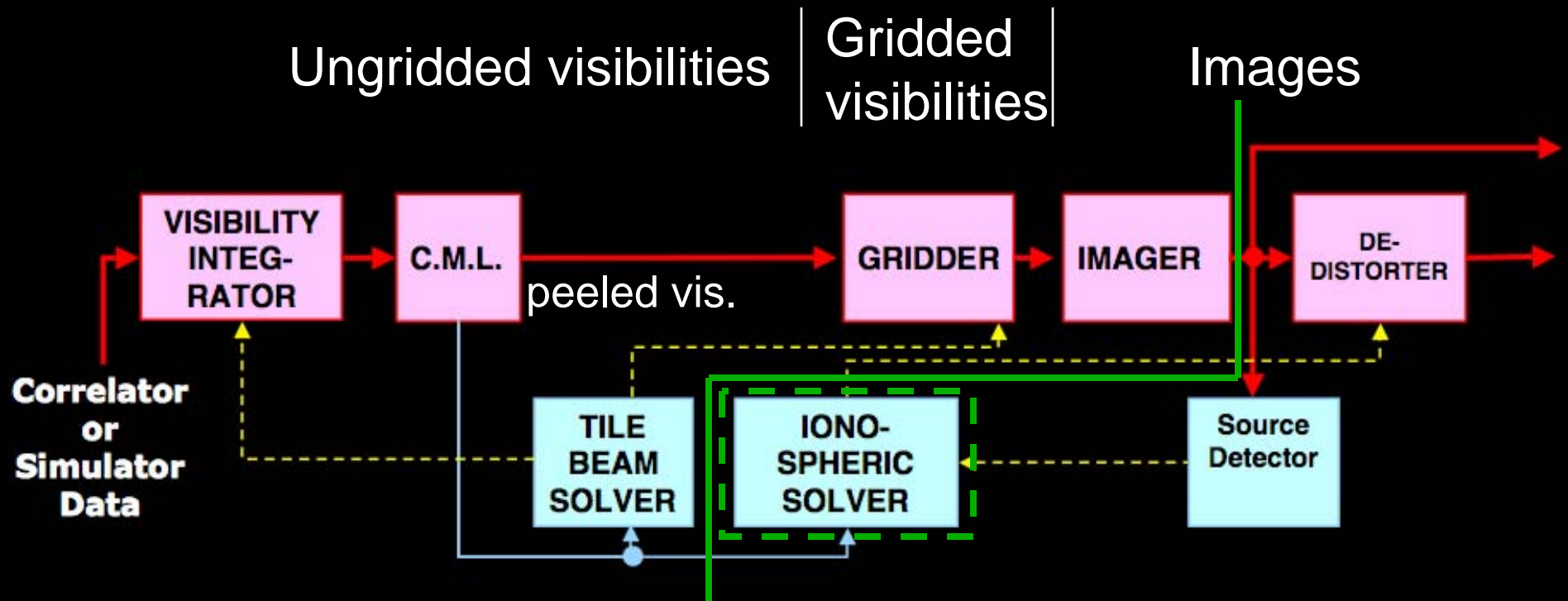


RTS Processing Flow



- Parallel-by-frequency processing steps; data sedentary w/in nodes
- Distribution of measurements (low-BW)
- Distribution of calibration (low-BW)

RTS Processing Flow



- Parallel-by-frequency processing steps; data sedentary w/in nodes
- Distribution of measurements (low-BW)
- → Distribution of calibration (low-BW)

RTS v.0 Elements in Harness

Element	TBD (examples)
<ul style="list-style-type: none">• Raw vis. integrator	<ul style="list-style-type: none">• σ^2-weighting
<ul style="list-style-type: none">• Cal Msrmt Loop	<ul style="list-style-type: none">• ν-chan avg'g; polarized sky; tolerance testing
<ul style="list-style-type: none">• Tile-beam solver	<ul style="list-style-type: none">• full beam response; tolerance testing;
<ul style="list-style-type: none">• Gridder	<ul style="list-style-type: none">• kernel (w-corr; fully integrate beam response; optimal anti-alias)
<ul style="list-style-type: none">• Imager	

RTS v.0 Elements to be Integrated

Element	TBD (examples)
<ul style="list-style-type: none">• Source Detector• Ionospheric solver:• De-distorter:	<ul style="list-style-type: none">• planning• trial integration• planning
<ul style="list-style-type: none">• Simulations (ongoing)	<ul style="list-style-type: none">• clean Stokes-I diffuse sky model; polarized sky model; translating turbulent ionosphere

Simple Tile Beam: 10%, 5° noise

QuickTime™ and a
GIF decompressor
are needed to see this picture.

QuickTime™ and a
GIF decompressor
are needed to see this picture.

CML Peel Demonstration

- **XX** polarization
- 140 MHz
- **Compact** sky only
- 10^3 brightest sources
- LST=1.5^h Boolardy
- Flat sky projection

QuickTime™ and a
GIF decompressor
are needed to see this picture.

CML Peel Demonstration

- **XX** polarization
- 140 MHz
- **Compact/diffuse** sky
- 10^3 brightest sources
- LST=1.5^h Boolardy
- Flat sky projection

QuickTime™ and a
GIF decompressor
are needed to see this picture.

CML Peel Demonstration

- **StokesQ** polarization
- 140 MHz
- **Compact** sky only
- 10^3 brightest sources
- LST=1.5^h Boolardy
- Flat sky projection

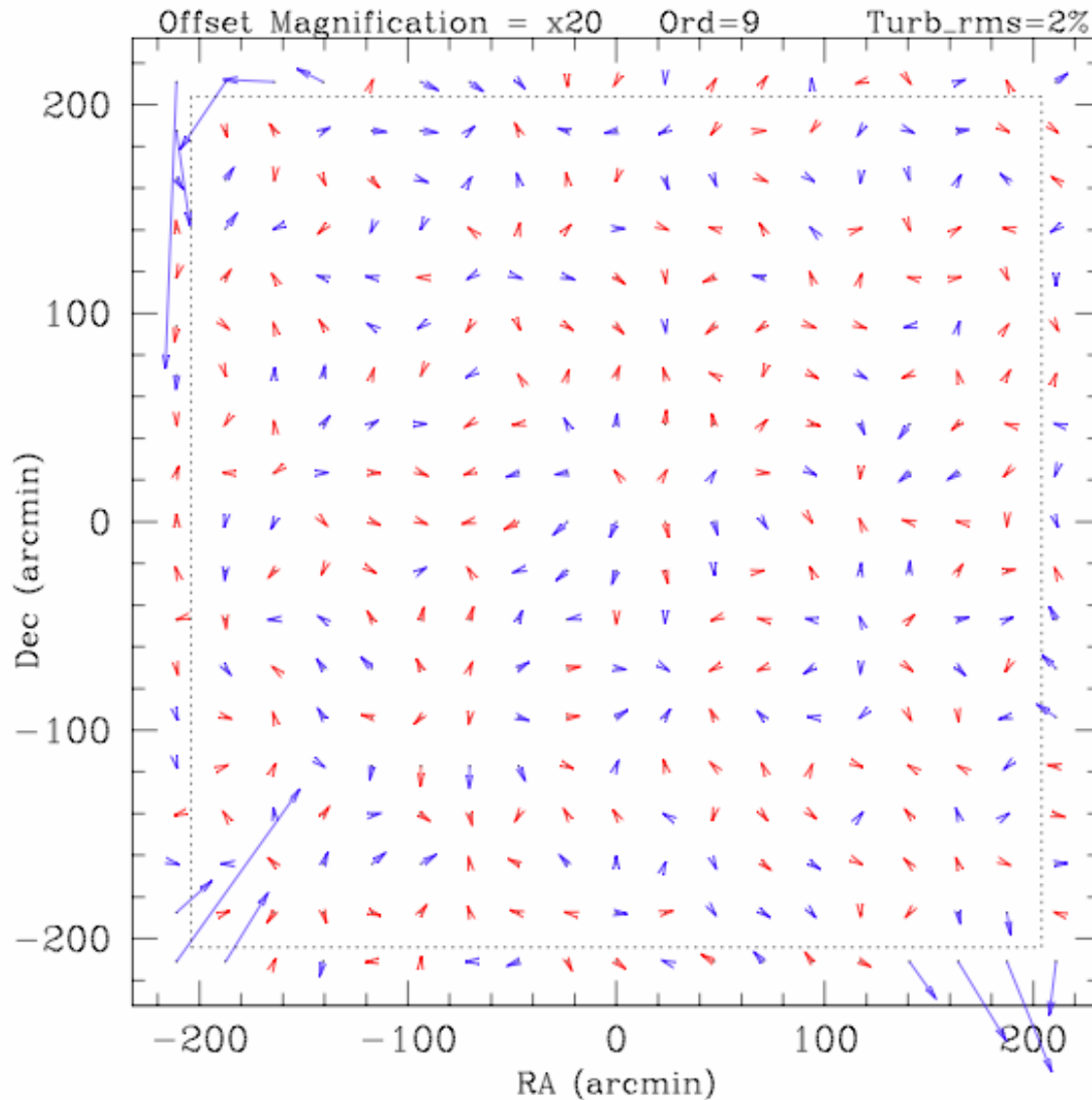
QuickTime™ and a
GIF decompressor
are needed to see this picture.

CML Peel Demonstration

- StokesQ polarization
- 140 MHz
- Compact/diffuse sky
- 10^3 brightest sources
- LST=1.5^h Boolardy
- Flat sky projection

QuickTime™ and a
GIF decompressor
are needed to see this picture.

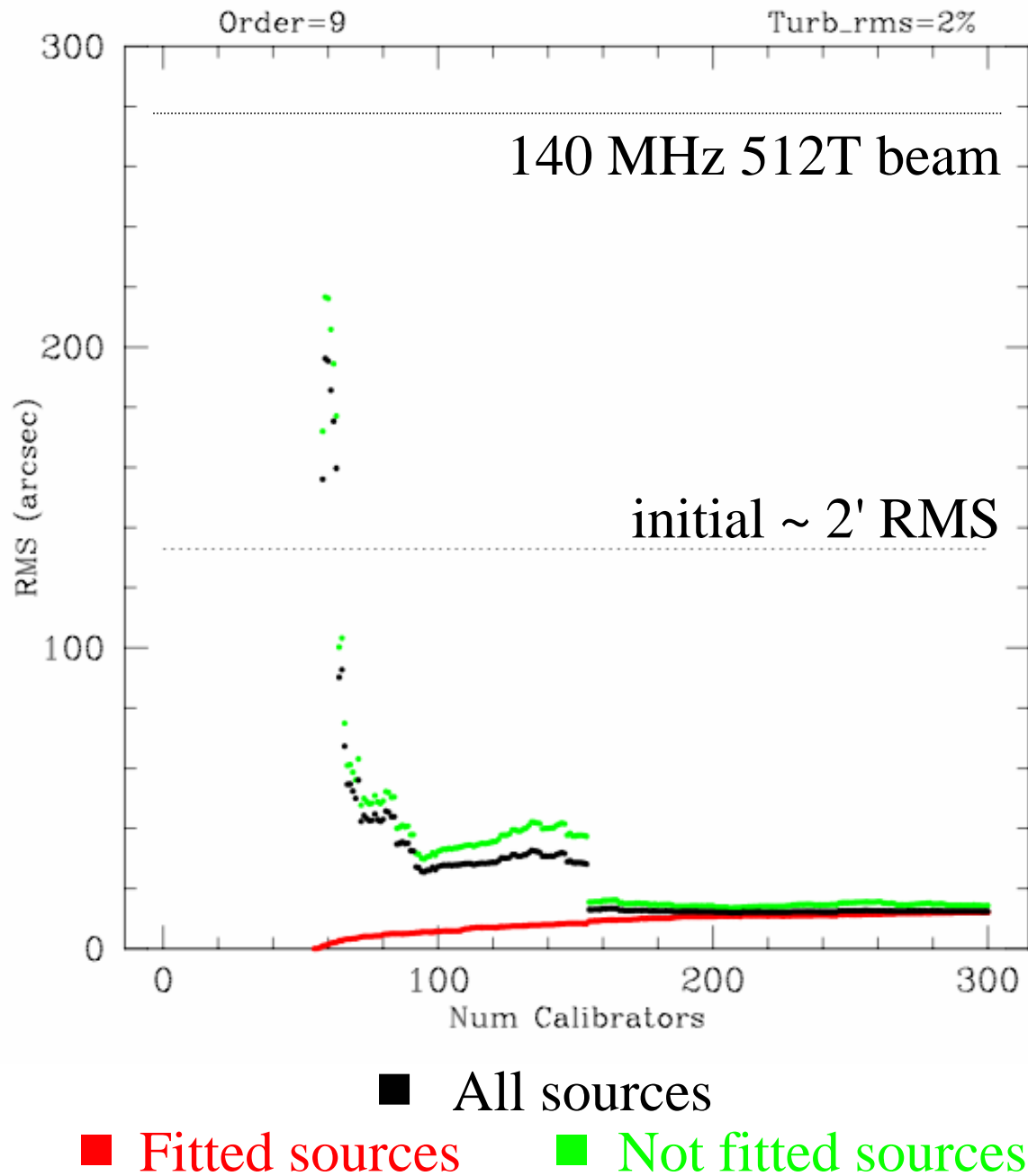
Ionospheric Correction Demo.



- Chapman profile
- 2% turbulence
 - ~ 4x VLA night
 - ~ 2x VLA day
- 9th order poly.
- 361 sources
- 8°x8° field; 140 MHz

→ Fitted sources → Sample for RMS est.

Ionospheric Correction Demo.



- Chapman profile
- 2% turbulence
 - ~ 4x VLA night
 - ~ 2x VLA day
- 9th order poly.
- 361 sources
- 8°x8° field; 140 MHz
- 30°x30° field will be a challenge.
 - Source Detector
 - >1000 sources req.

End-to-end Processing

- Sky to MAPS sim
MAPS to RTS
RTS to display
- Demo CML
- Demo Gridder
- Demo data path
 - Compact/diff. sky
 - 10^3 brightest srcs
 - 20 sources peeled
 - XX @ 140 MHz
 - LST=6^h Boolardy

QuickTime™ and a
TIFF (LZW) decompressor
are needed to see this picture.

36° square

Relevance to 32T

- **32T enables RTS** real-world testing & guides development
 - benchmarks inform scaling of computation to 512T
- **RTS enables 32T** data processing **off-site**
 - Assessment of dipole/tile performance
 - Construction of initial un/polarized sky survey
- **RTS testing transitions to site with h/w transition to 512T**
 - regulated by RTC/infrastructure development

CML Peel 32T Demonstration

- **XX** polarization
- 140 MHz
- **Compact** sky only
- 10^3 brightest sources
- LST=1.5^h Boolardy
- Flat sky projection

QuickTime™ and a
GIF decompressor
are needed to see this picture.

CML Peel 32T Demonstration

- **XX** polarization
- 140 MHz
- **Compact/diffuse** sky
- 10^3 brightest sources
- LST=1.5^h Boolardy
- Flat sky projection

QuickTime™ and a
GIF decompressor
are needed to see this picture.

end

Analytic Tile Beam: no noise

